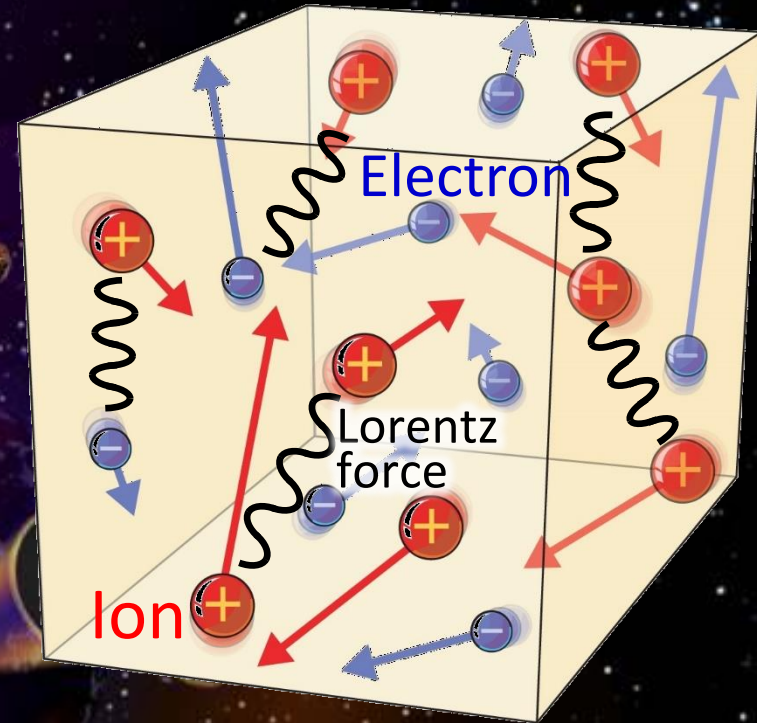




# Computer Simulations on Rocket-Plasma Interactions

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# Space plasma environment



Space filled with an ionized gas called “plasma”

- consisting of huge number of free-moving electrons and ions
- interactions with electromagnetic field
- interactions with rocket or spacecraft

# Basic equations

## Charged particle dynamics & electromagnetic evolution



Coupled

Newton's equations  
of particle motion

$$\frac{dv_i}{dt} = \frac{q_s}{m_s} (\mathbf{E} + \mathbf{v}_i \times \mathbf{B})$$
$$\frac{dx_i}{dt} = \mathbf{v}_i$$

Field equations

$$\nabla \times \mathbf{E} = -\frac{\partial \mathbf{B}}{\partial t}$$

Full-electro-magnetic treatment

$$\nabla \times \mathbf{B} = \mu_0 \mathbf{J} + \frac{1}{c^2} \frac{\partial \mathbf{E}}{\partial t}$$

or

$$\nabla \cdot \mathbf{E} = \frac{\rho}{\epsilon_0}$$

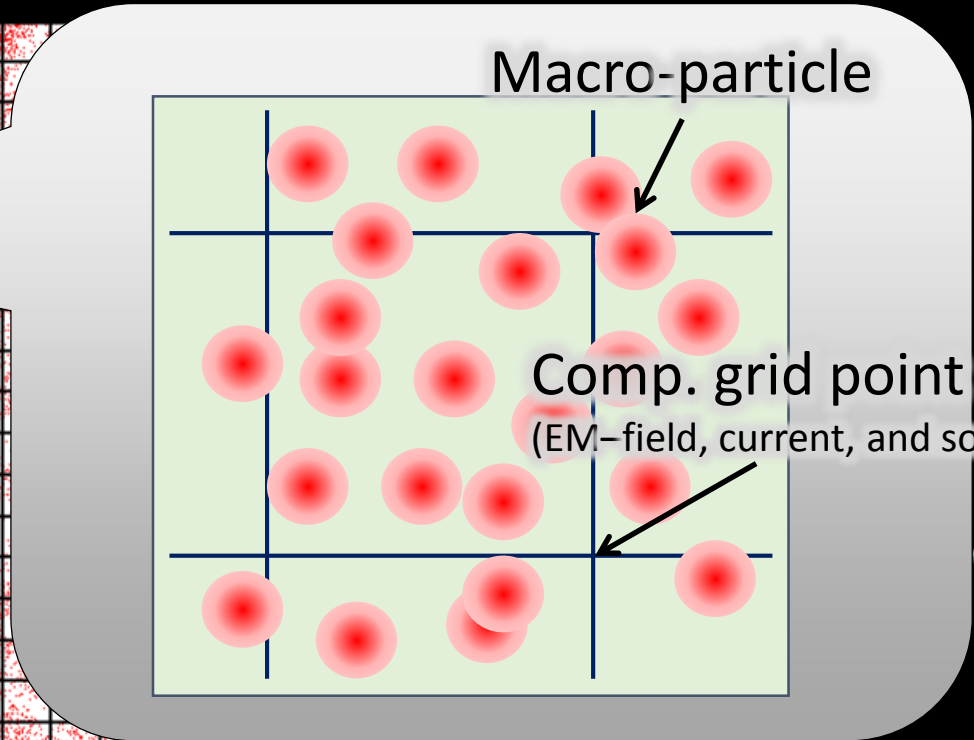
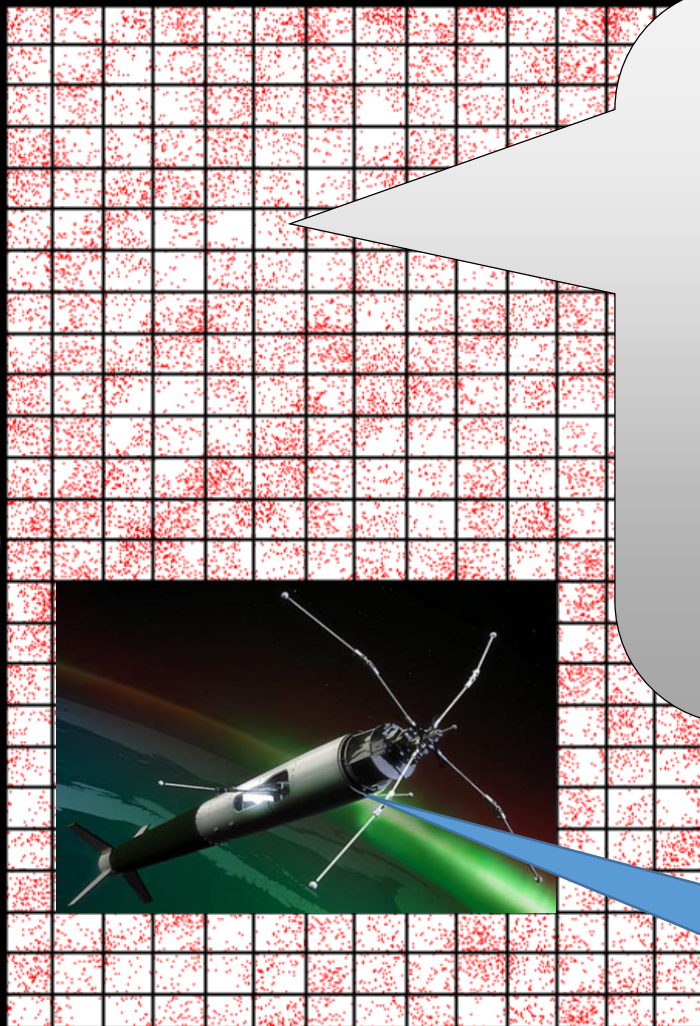
Electrostatic approximation

$$\nabla \cdot \mathbf{B} = 0$$

These two sets of basic eq. are simultaneously solved.

# Plasma particle simulation based on **particle-in-cell (PIC)** method

Simulation  
code "EMSES"



Rocket body described  
as internal boundary

# Necessity of Supercomputing

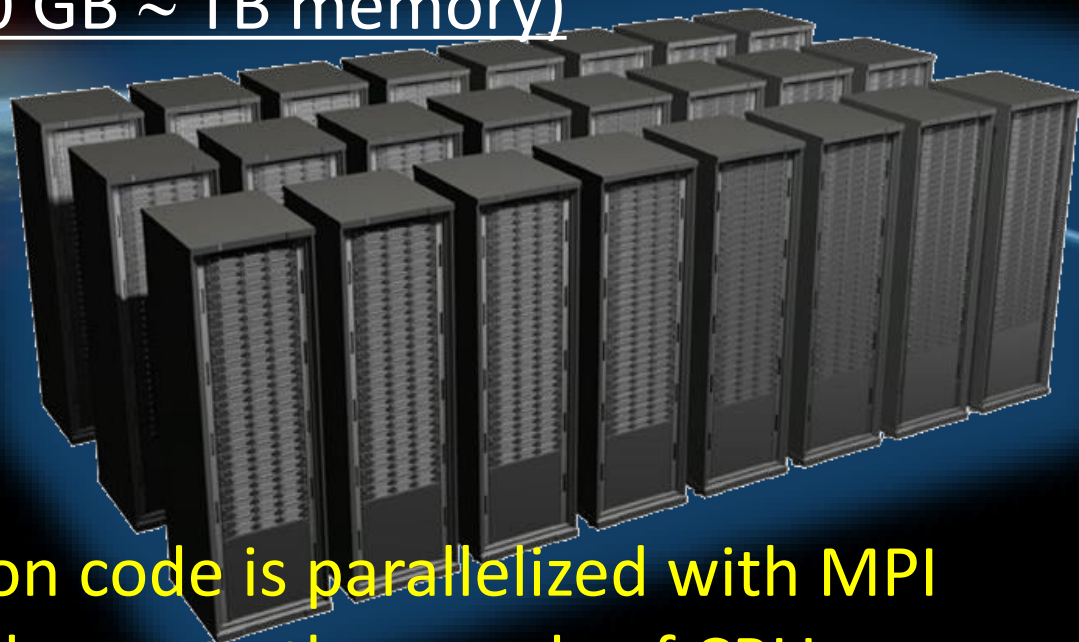
Debye length  $\lambda_D$ : smallest spatial scale in space plasma

System size of practical problems:  $100 \sim 1000 \lambda_D$

...need to handle

Total # of grids:  $10^6 \sim 10^8$  & # of particles:  $10^8 \sim 10^{10}$

(corresponding to 100 GB  $\sim$  TB memory)

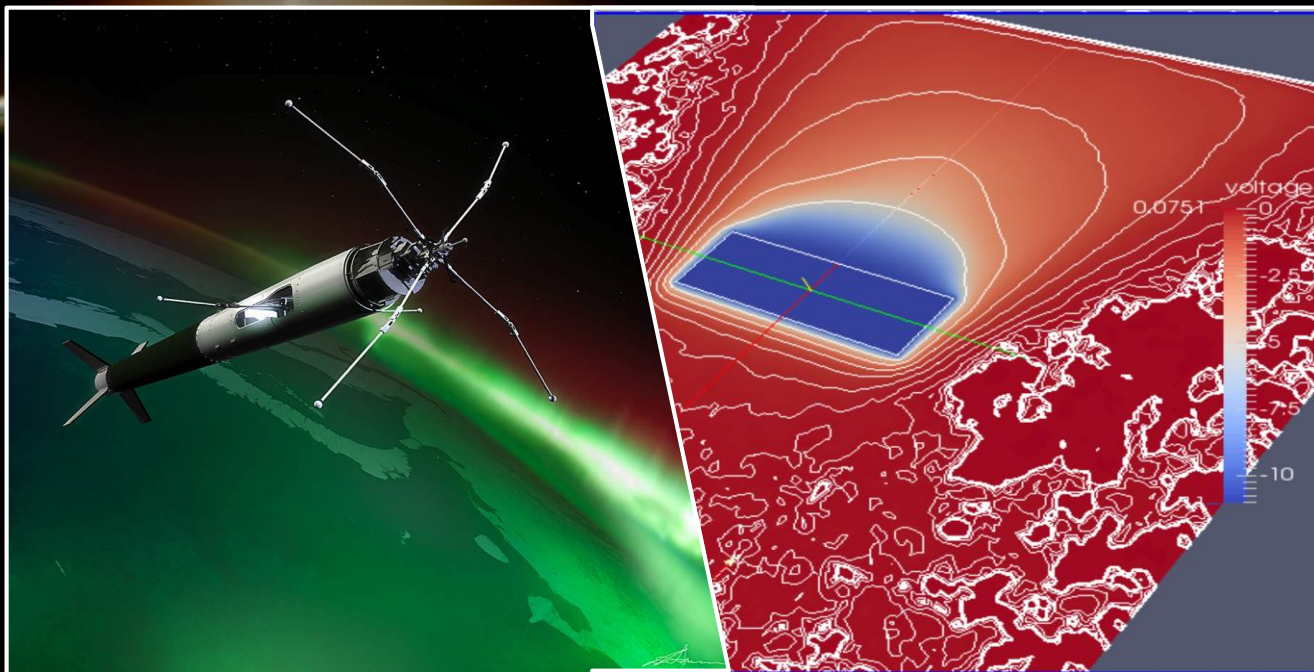


Our plasma simulation code is parallelized with MPI (& OpenMP) and scales up to thousands of CPU cores.



# Our goals...

To study interactions between the sounding rocket and the surrounding ionospheric plasma by means of numerical simulations using supercomputers, data analysis, and scientific visualization.



Ben



“Development & parallelization of new code”  
... based on essence of PIC method

Alexis



“Simulation results on rocket-plasma interactions”  
... presented with fascinating 3D movies

Takuya



“Focused analysis on plasma particle dynamics”  
... going inside details of physics

I'm so satisfied with their works!



# Interactions between sounding rocket & ionospheric plasma



We focus on **rocket experiments to investigate ionospheric plasma environment.**

→ Understanding of rocket-plasma interactions is essential to operate such rocket experiments appropriately.

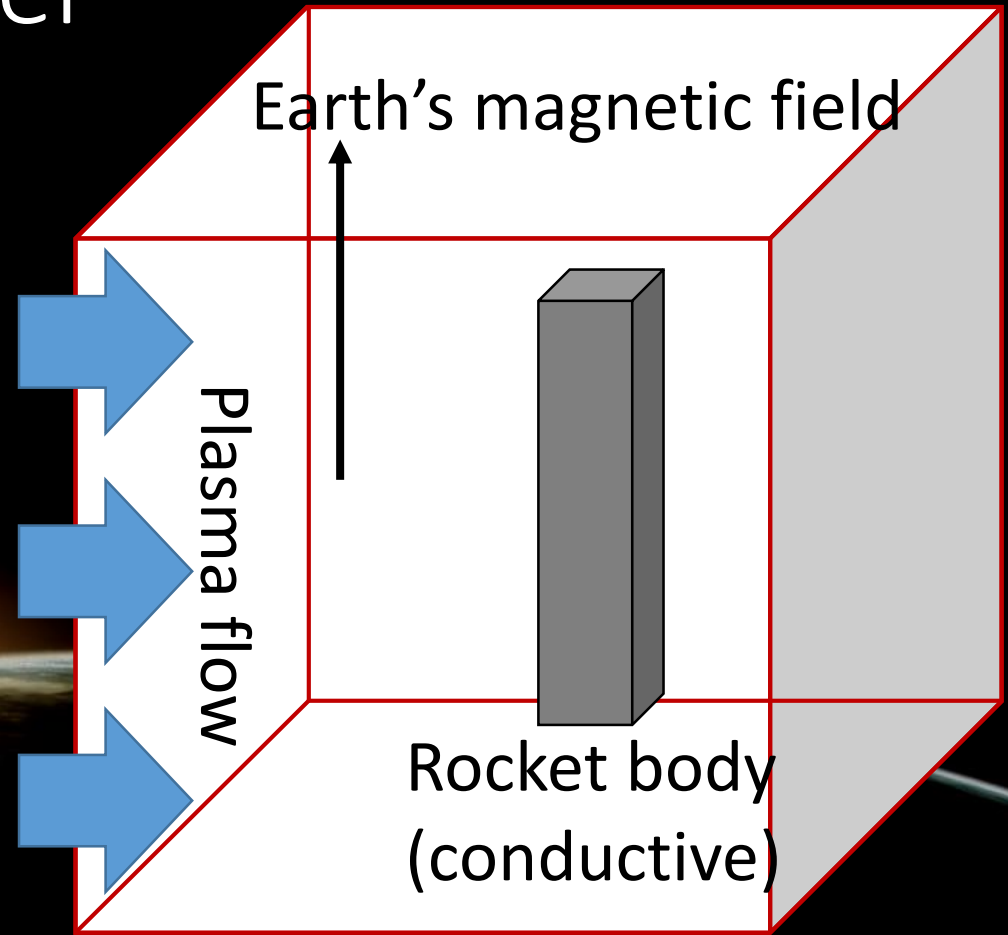
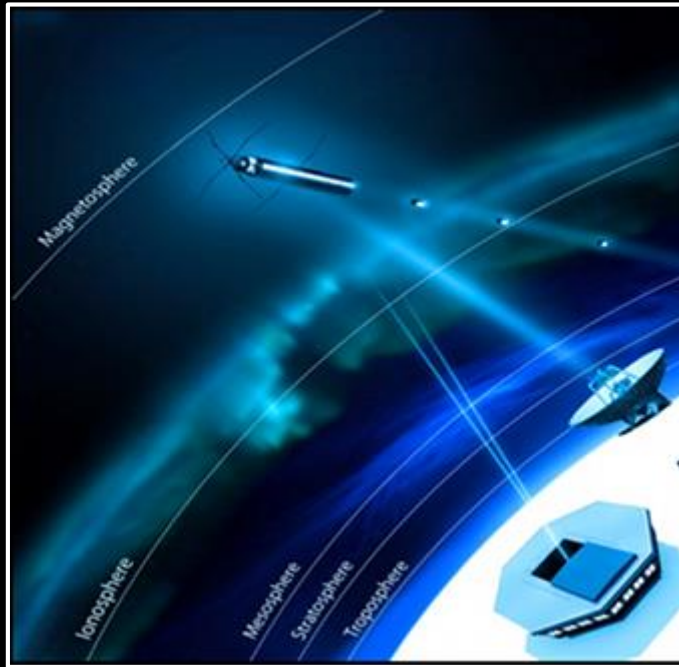
→ **Aid of numerical simulations**



# What to see?

1. How does the rocket body charge?
2. What is the plasma (electron & ion) distributions around the rocket?
3. What is the plasma dynamics around the rocket? Is it possible to interpret the dynamics from an electric potential profile (& Earth's magnetic field) around the rocket?

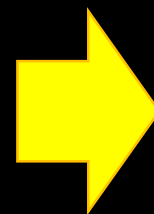
# Simulation model



## Problem size

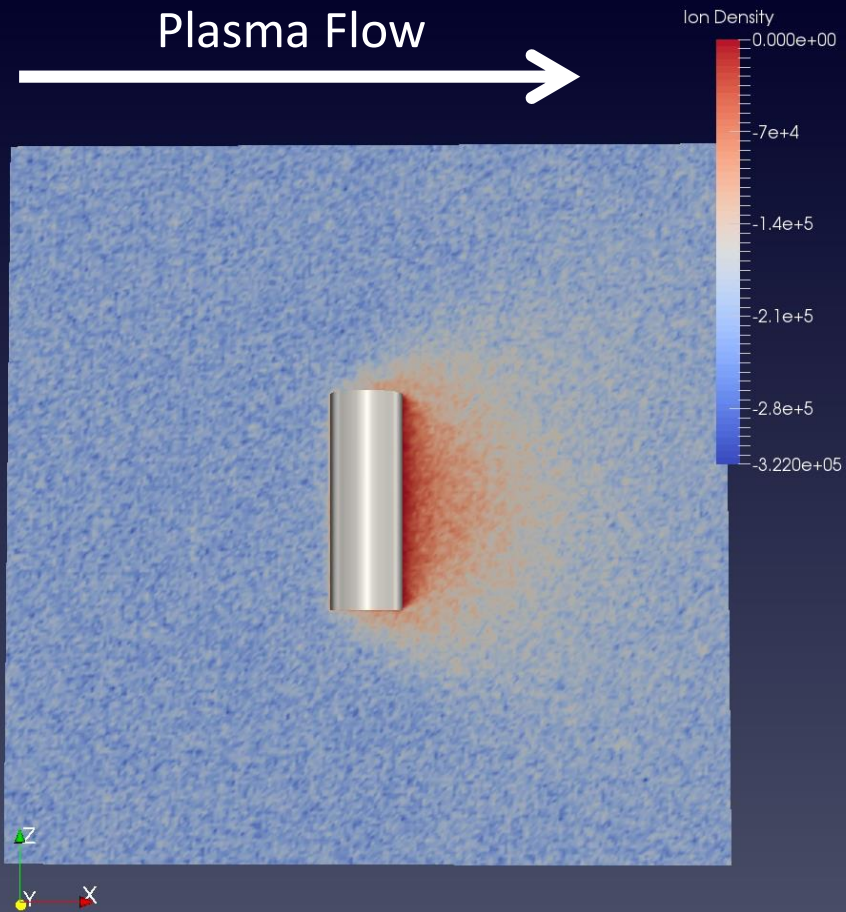
- # of grids:  $240 \times 240 \times 240$
- # of particles:  $\sim 10^9$
- ...requires  $>1$  TB memory

## Simulation domain

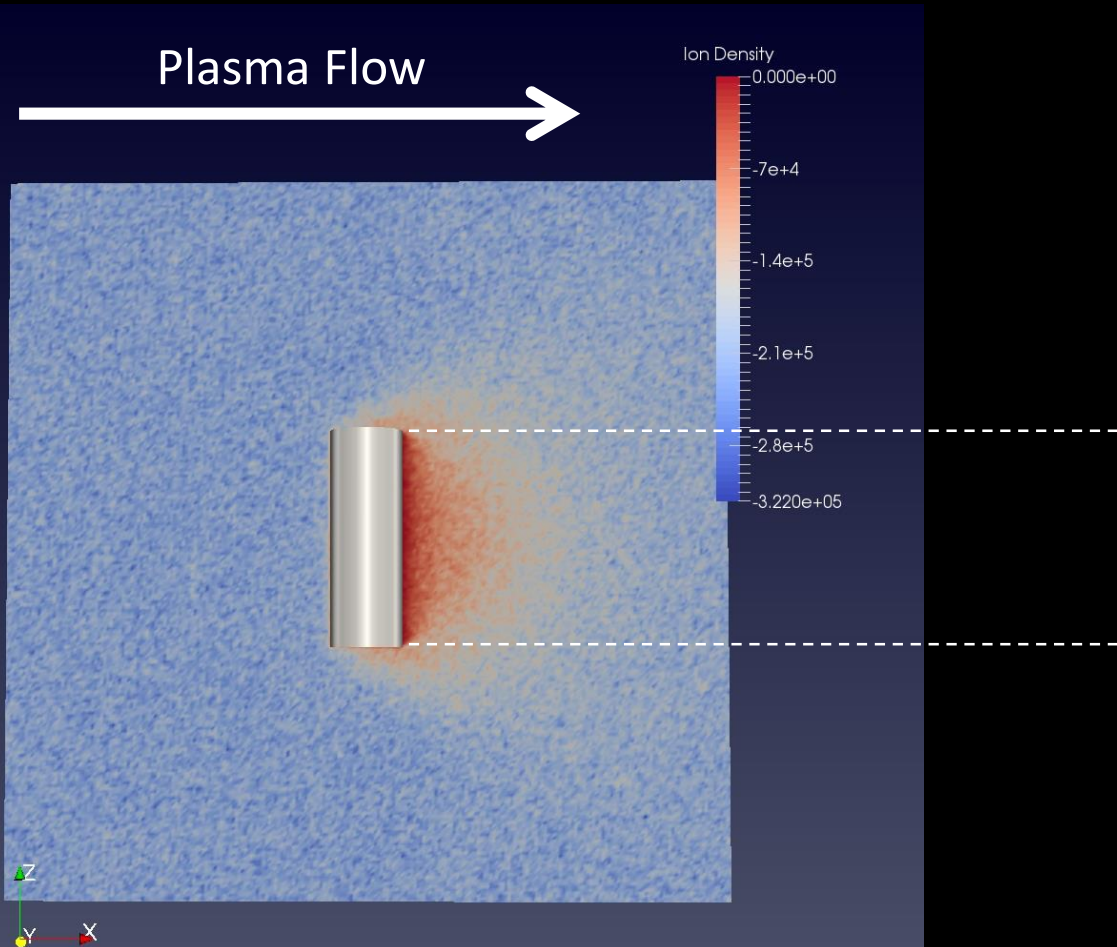


Use 384 CPU-cores of Kobe FX10 supercomputer

# Particle Densities

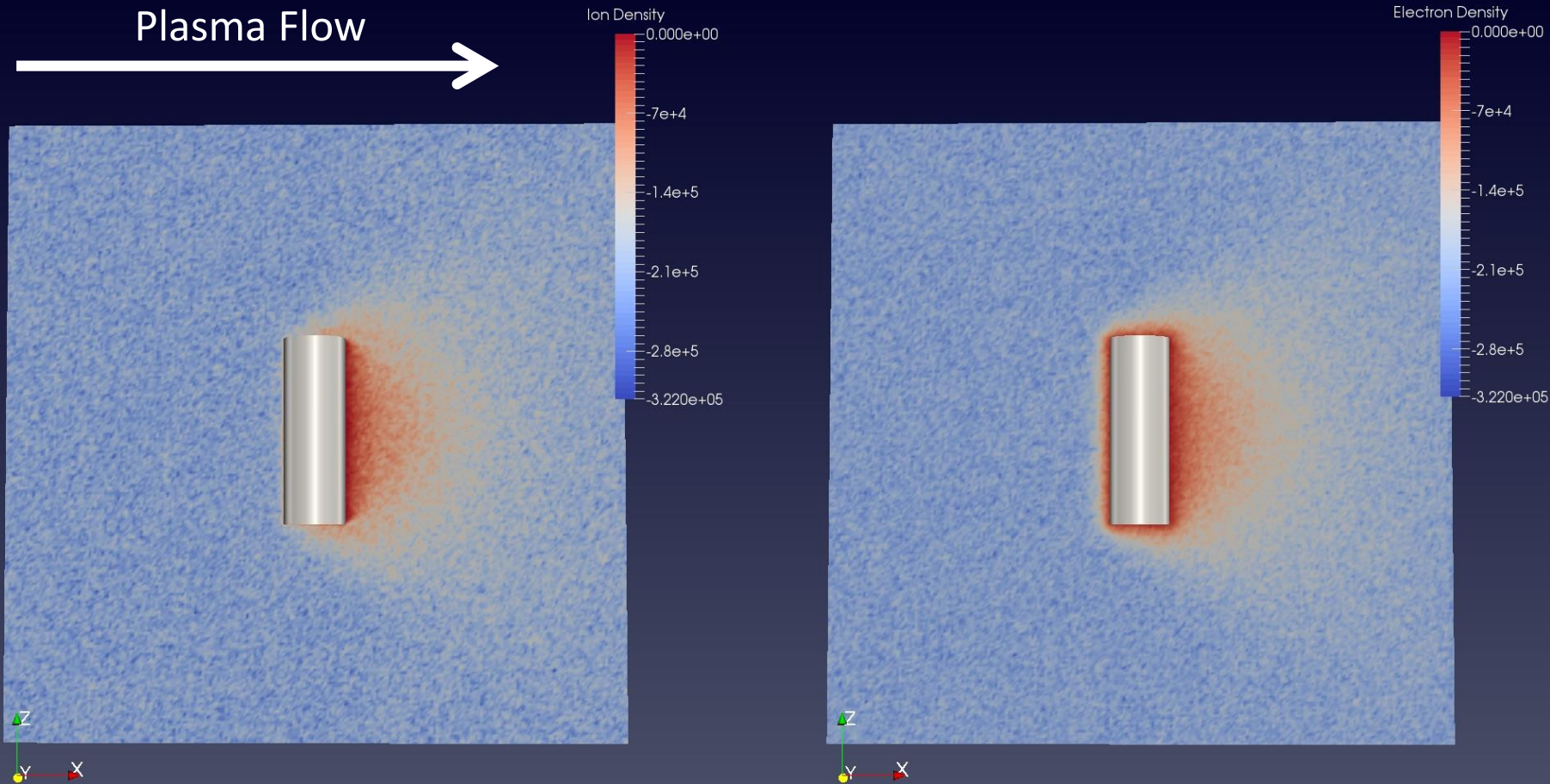


# Particle Densities





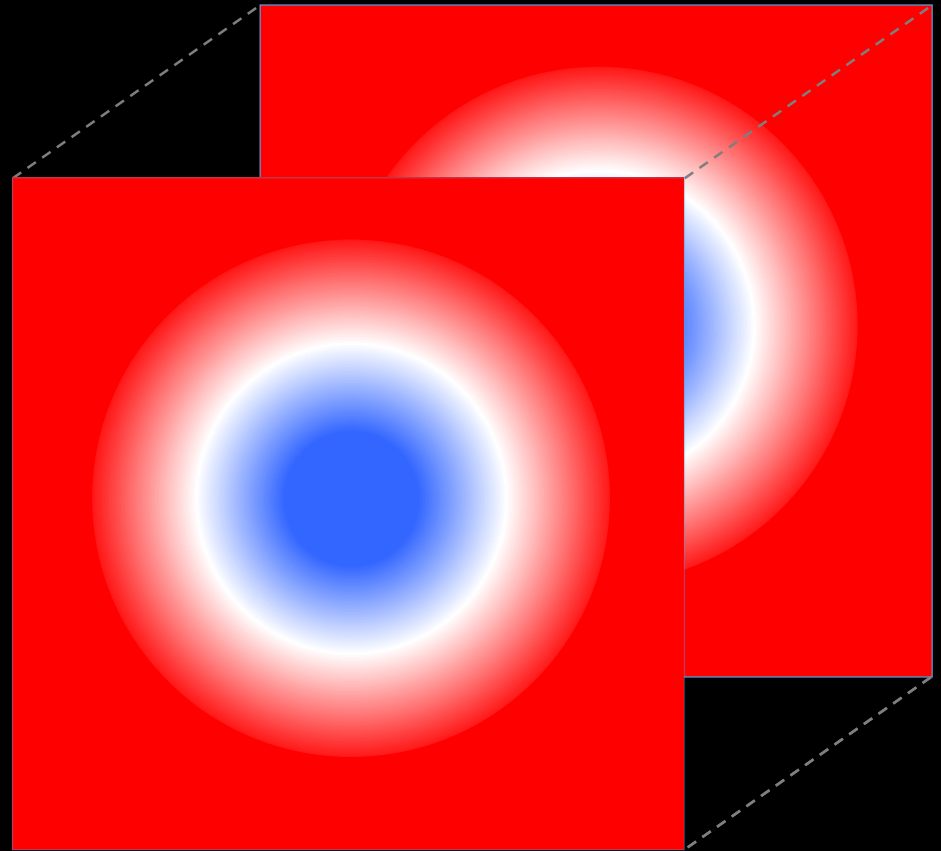
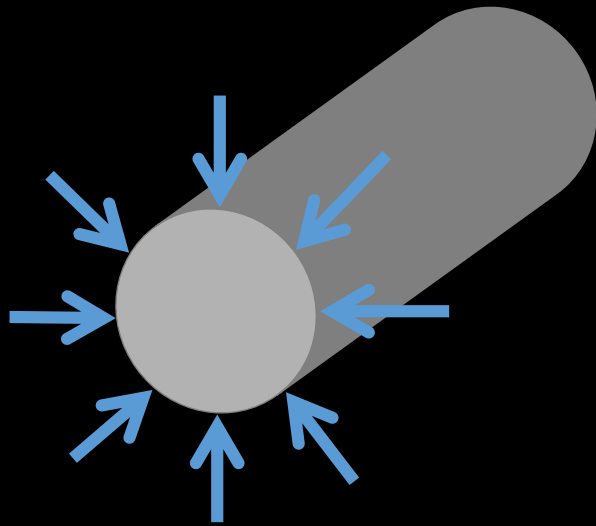
# Particle Densities



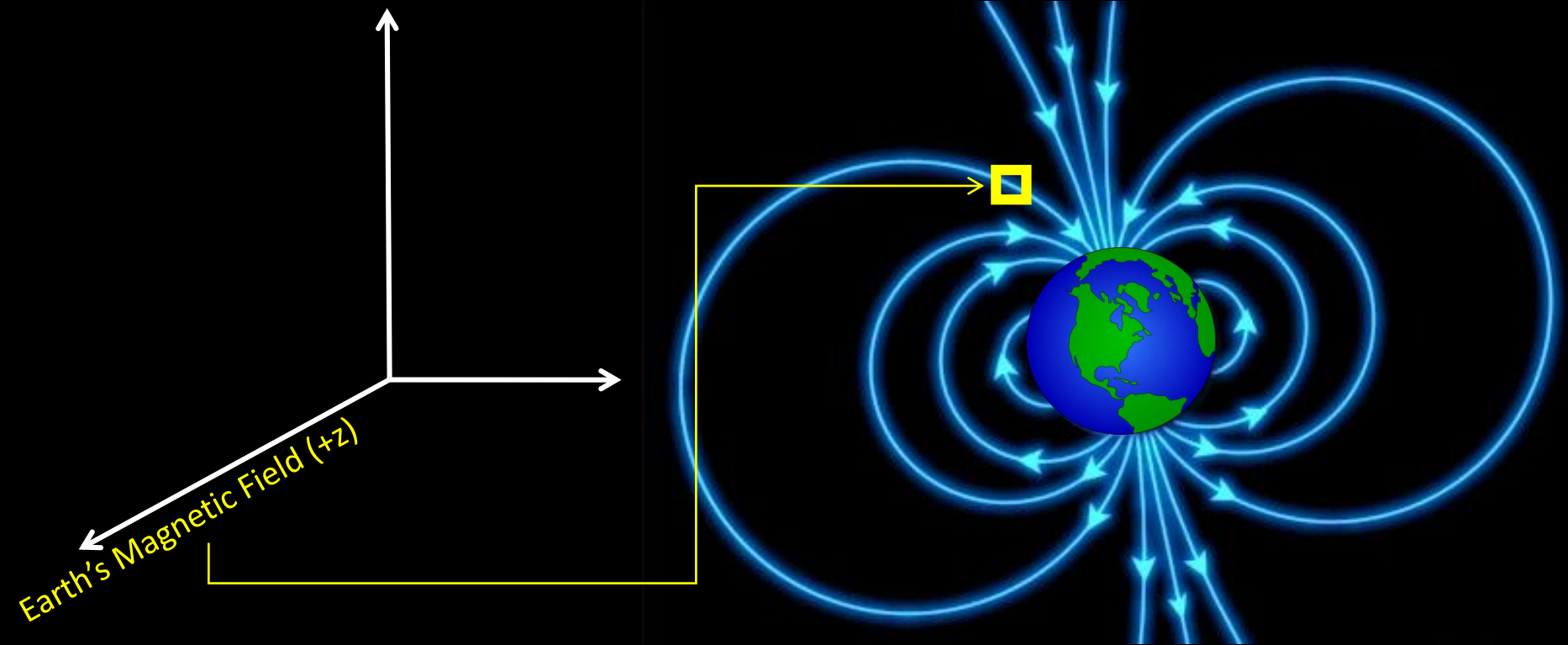
Play movie here

3D movie

# Rocket Potential

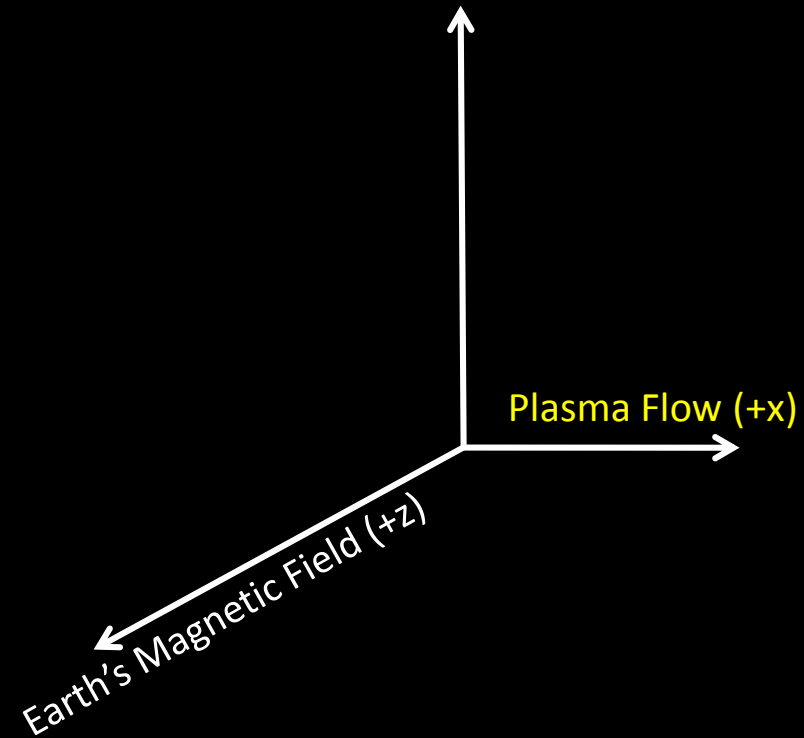


# Plasma Potential

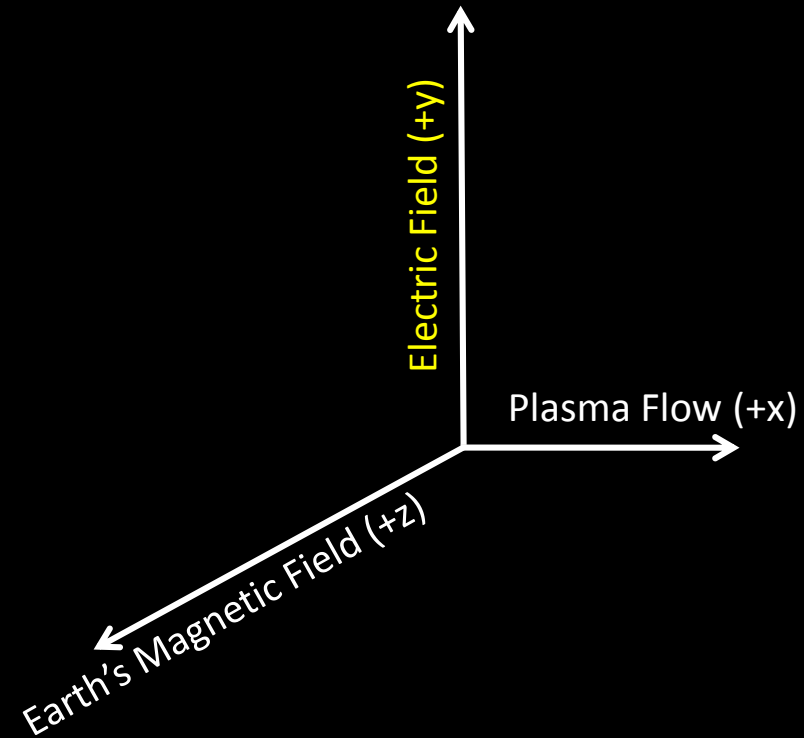




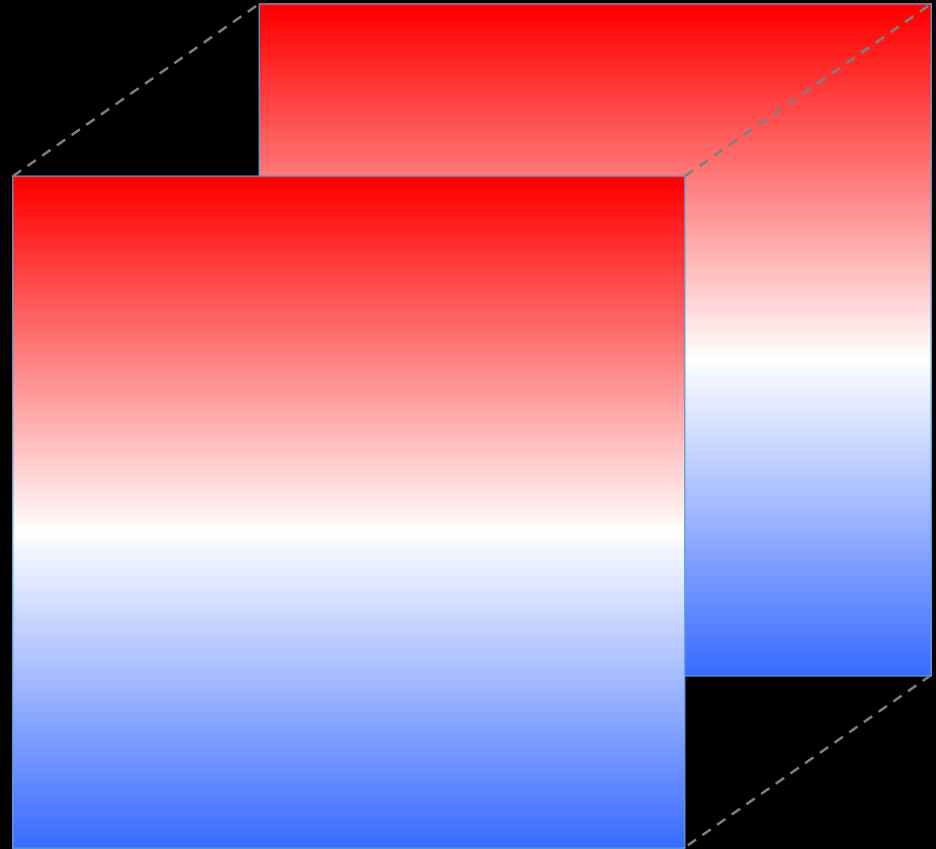
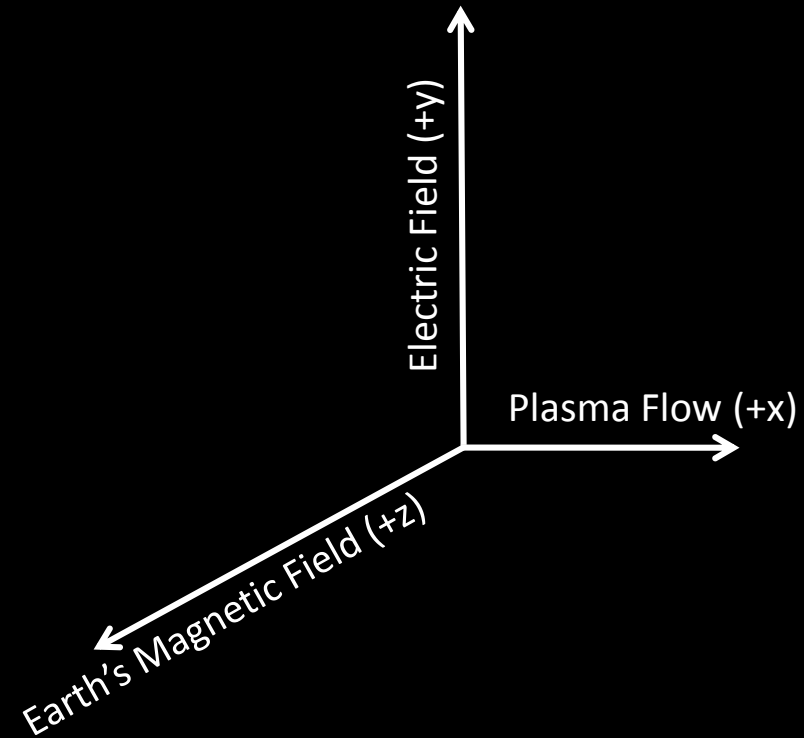
# Plasma Potential



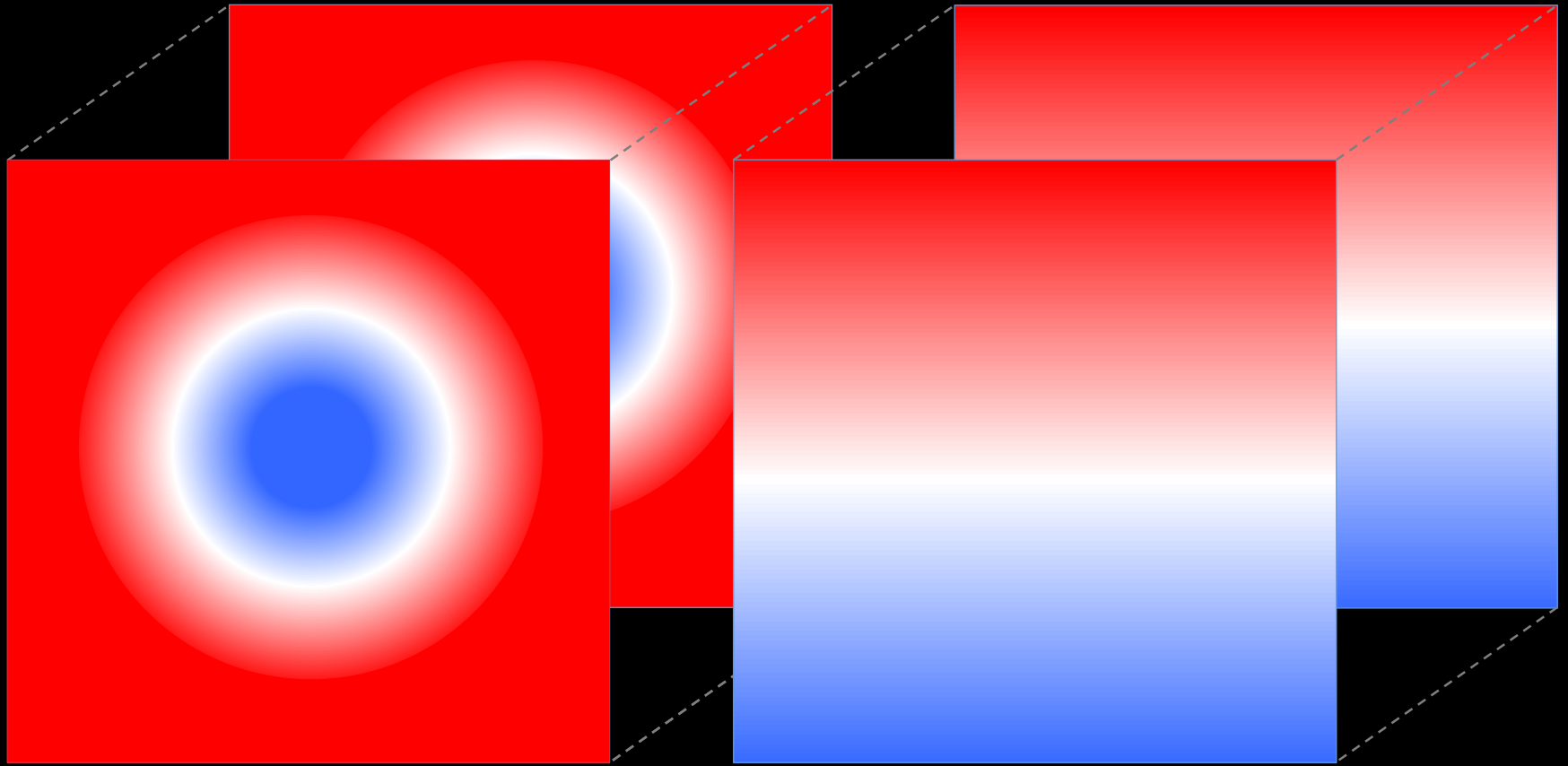
# Plasma Potential



# Plasma Potential



# EMSES Potential

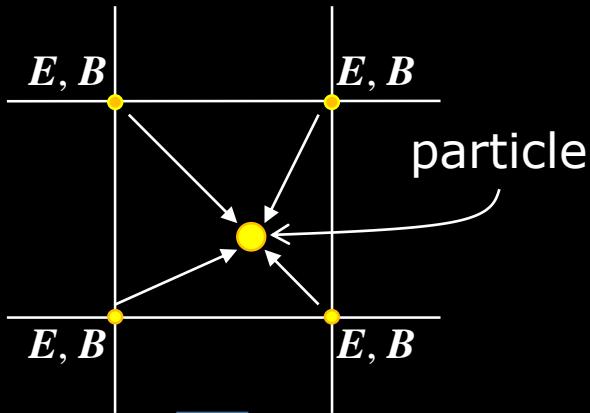




Play movie here

3D movie

# EMSES plasma simulation: main loop



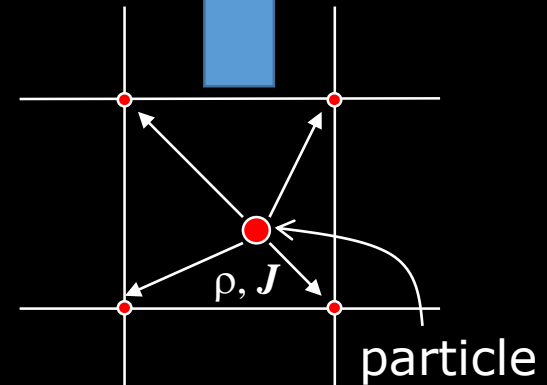
$$\nabla \times \mathbf{E} = -\frac{\partial \mathbf{B}}{\partial t}$$

$$\nabla \times \mathbf{B} = \frac{1}{c^2} \frac{\partial \mathbf{E}}{\partial t} + \mu_0 \mathbf{J}$$

or

$$\nabla \cdot (-\nabla \phi) = \frac{\rho}{\epsilon_0}$$

Field evolution



Current/charge deposition

$$\frac{d\mathbf{v}}{dt} = \frac{q_s}{m_s} (\mathbf{E} + \mathbf{v} \times \mathbf{B})$$

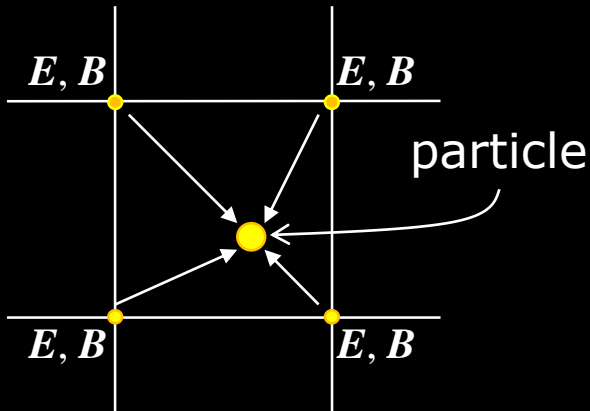
$$\frac{d\mathbf{r}}{dt} = \mathbf{v}$$

Particle dynamics

$\Delta t$

# Development of test particle code

Using the field profile obtained from EMSES simulation



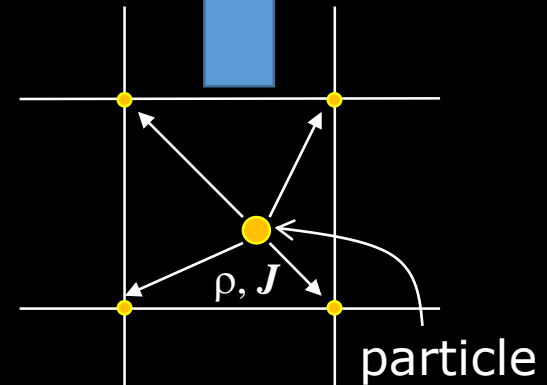
~~$$\nabla \times \mathbf{B} = \frac{1}{c} \frac{\partial \mathbf{E}}{\partial t} + \mu_0 \mathbf{J}$$

or

$$\nabla \cdot (-\nabla \phi) = -\frac{\rho}{\epsilon_0}$$~~

Field evolution

$\Delta t$



Current/charge deposition

$$\frac{d\mathbf{v}}{dt} = \frac{q_s}{m_s} (\mathbf{E} + \mathbf{v} \times \mathbf{B})$$

$$\frac{d\mathbf{r}}{dt} = \mathbf{v}$$

Particle dynamics

# Argument for Test Particle Code

EMSES simultaneously solves Maxwell's equations and Newton's laws (field evolution and particle motion), so we can already find where the particles are. Why should we write separate test particle code?

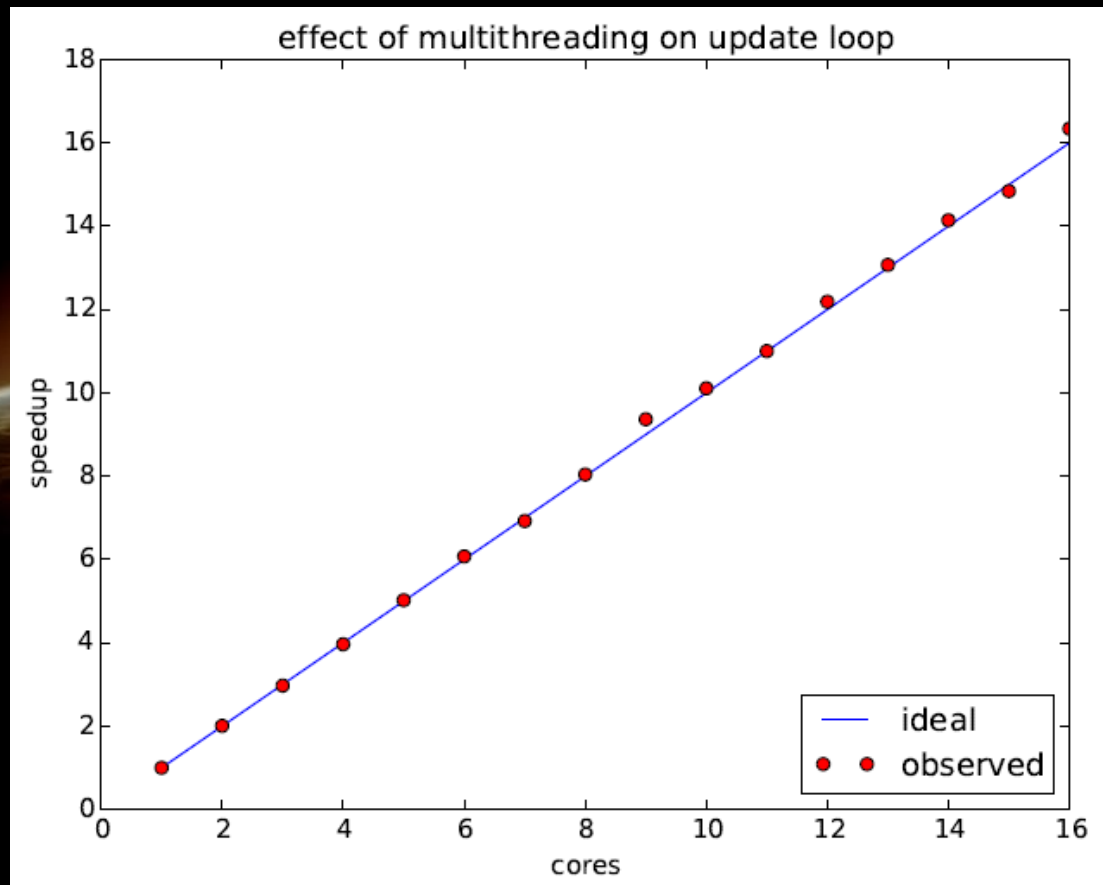
- interest in steady-state dynamics
- particles evolve according to decoupled system of ODEs (opportunity for parallelization)
- data locality and memory access

# Storage of $\vec{E}$ and $\vec{B}$

- In EMSES,  $E_x$ ,  $E_y$ , and  $E_z$  are stored as separate arrays (and the same for  $\vec{B}$ ).
- If we only care about test particles, is this ideal?
- Upfront cost of restructuring storage in exchange for potential benefits during long-running computation. Profiling required.

# Parallelization

Distribute particles among cores using OpenMP. Results with  $2^{20}$  particles and  $10^3$  timesteps:



Takes around 24 seconds with 16 cores



# Pop Quiz

Consider the following Fortran program.

```
call CPU_TIME(t0)
do i = 1, n
    u = cross_function(u, v)
end do
call CPU_TIME(t1); print *, t1 - t0
call CPU_TIME(t0)
do i = 1, n
    call cross_subroutine(u, v, w)
    u = w
end do
call CPU_TIME(t1); print *, t1 - t0
```

Similar code is used in updating particle velocities.  
Which loop will be faster?

# Answers

Compiled with gfortran -O3 on my laptop and with mpifrtpx -O3 -Kfast on the Kobe FX10.

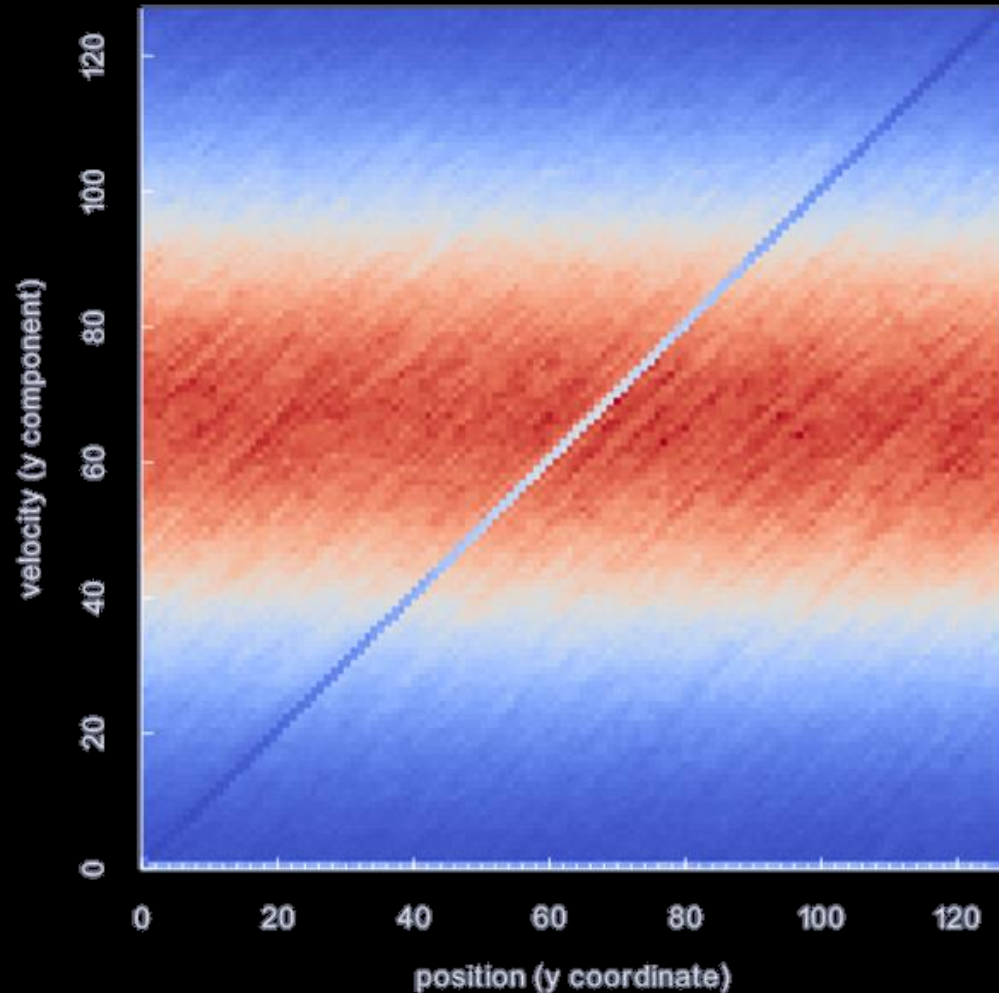
Results:

machine	function	subroutine
laptop	$1.2e-1$ s	$1.3e-1$ s
K computer	$6.2e+0$ s	$2.1e-1$ s

About **30 times** slower as a function on the Kobe FX10!

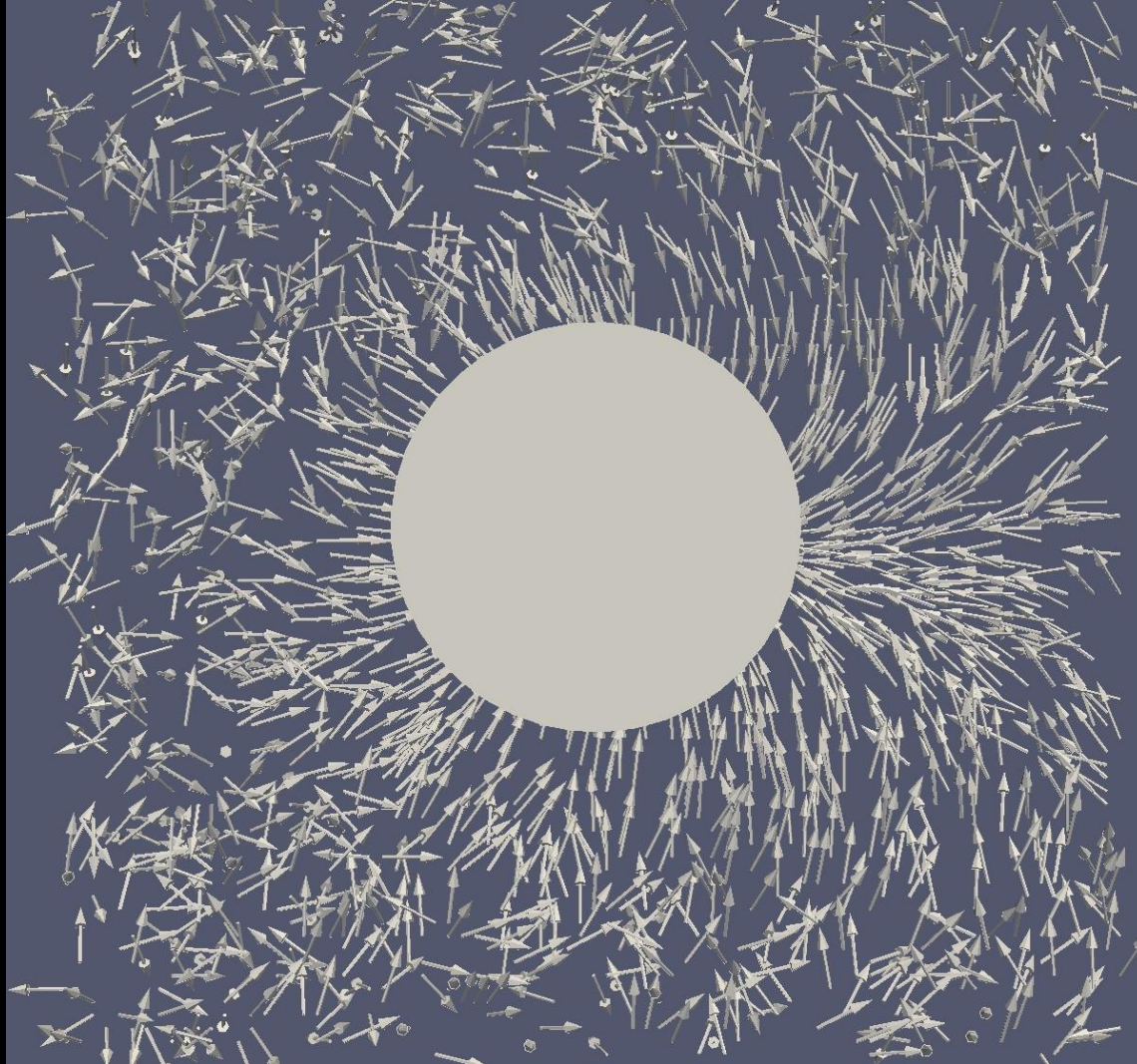
# Analytics

Now that we have the particle dynamics, what next?



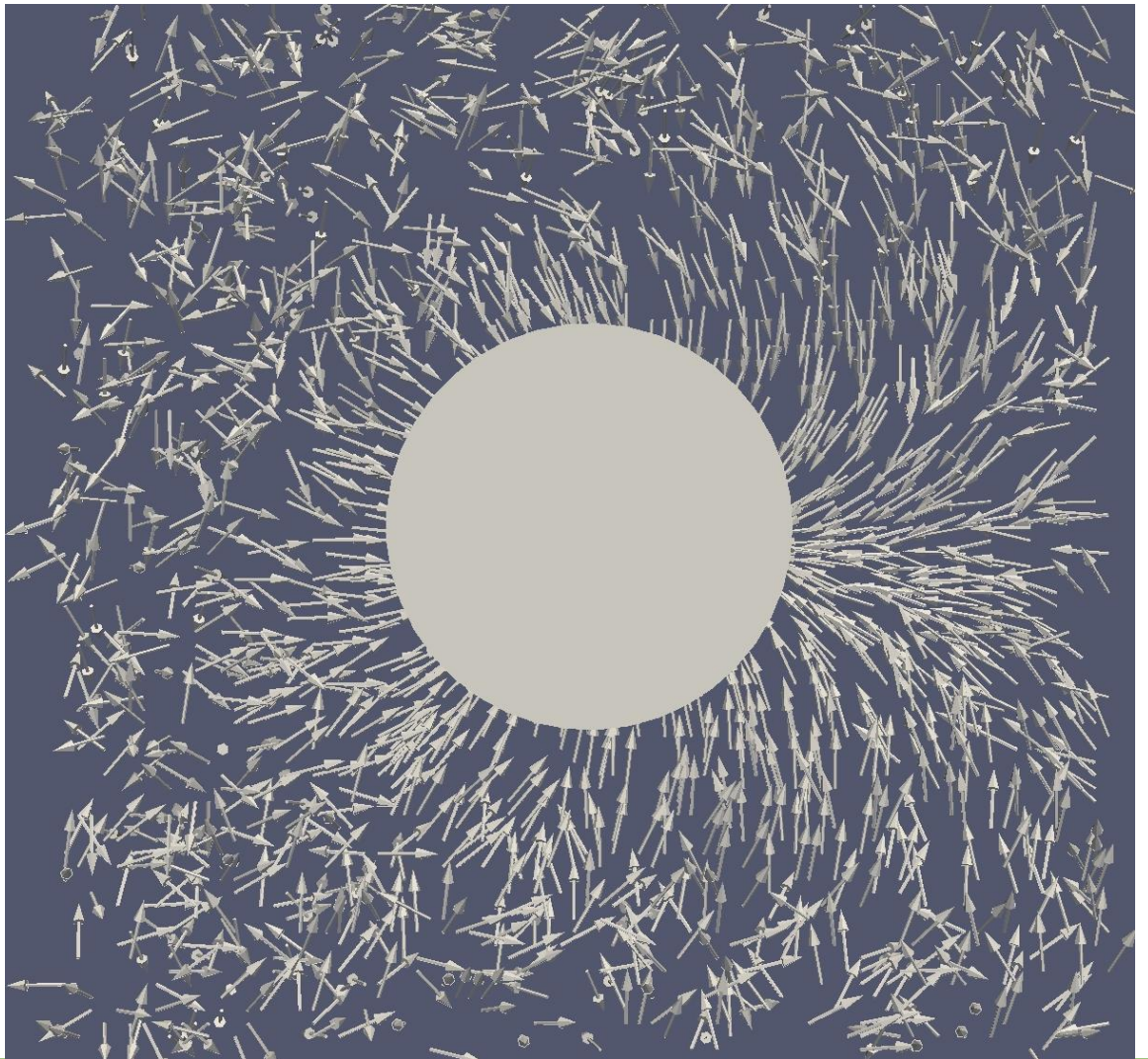
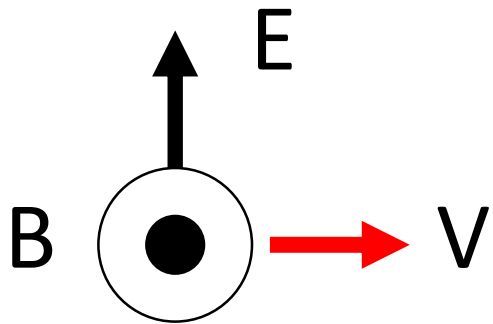
More opportunities for parallelization!

# Electric field from EMSES simulation



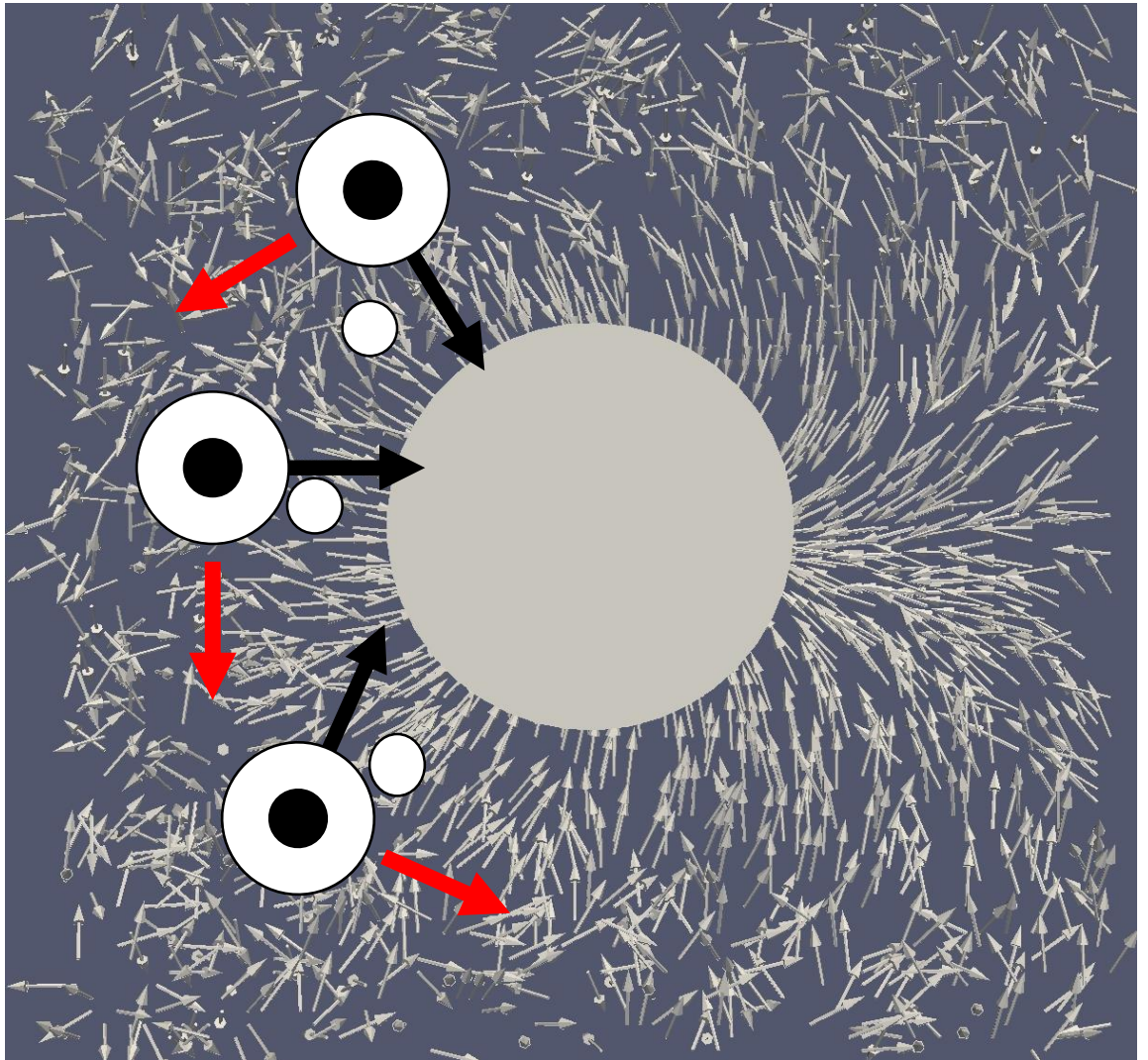


# Principle



# Principle

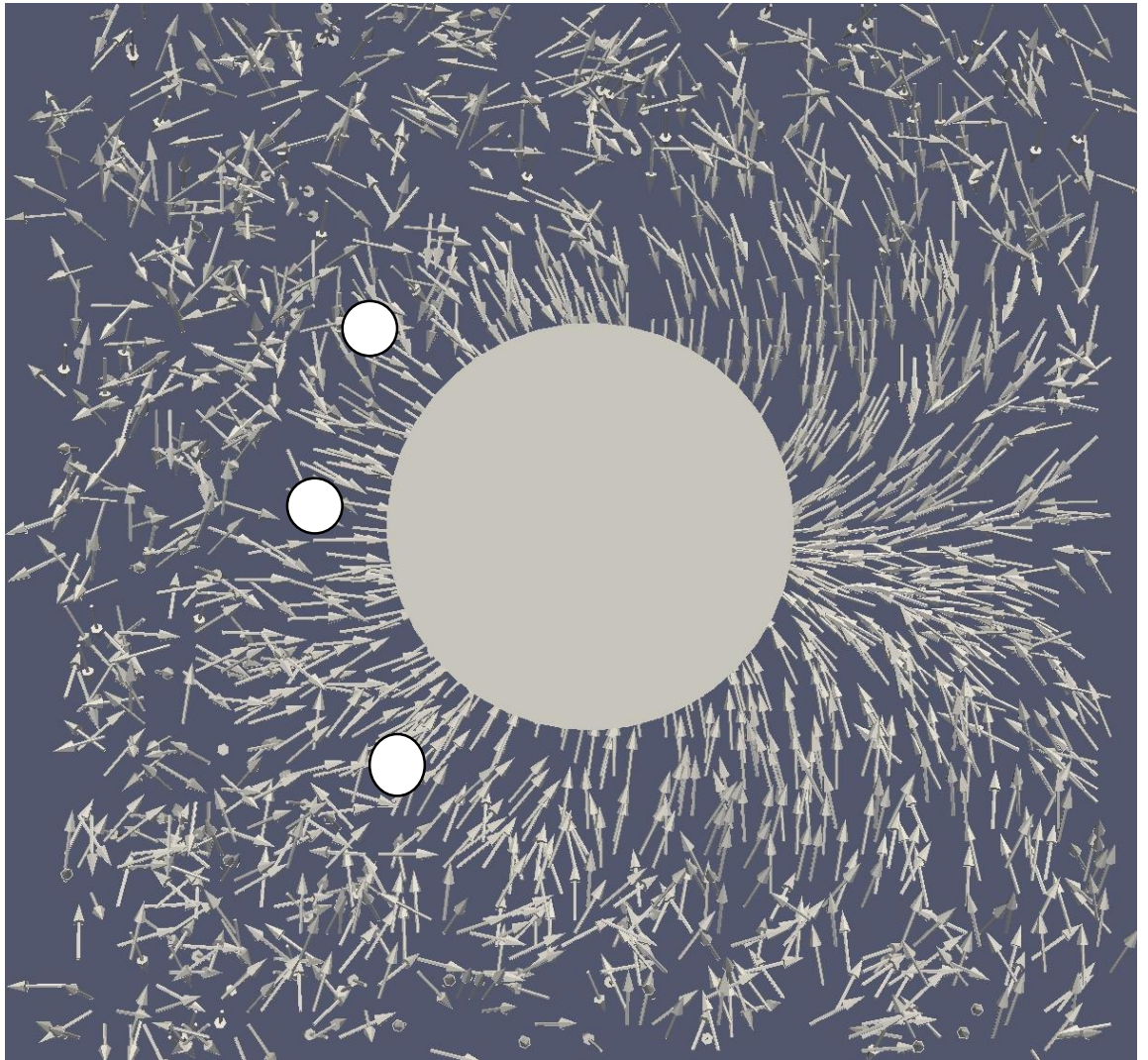
Direction  
Change of the  
electric field





# Principle

Counter  
Clockwise  
Rotation



# Summary

- We studied rocket plasma interactions by means of plasma particle simulations, data analysis and 3D visualization.
- We newly developed test particle code and parallelized the code with OpenMP.
- We confirmed rocket charging and plasma perturbations, and also understood characteristic particle dynamics due to a local electric field around the rocket.



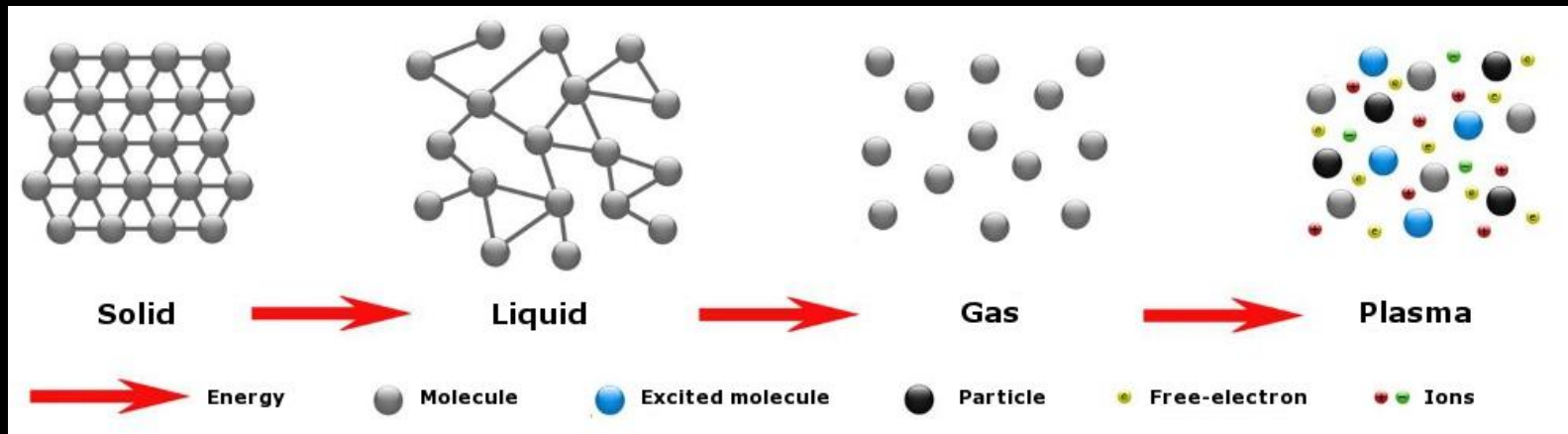
A vibrant nebula with pink and blue hues against a starry background. The nebula is the central focus, with a mix of soft pink and bright blue colors. It is surrounded by a dense field of stars, many of which are also blue or cyan in color. The overall scene is a rich, colorful representation of a star-forming region in space.

Thank you

Backup slides



# What is plasma?

Plasma is “the fourth state of matter.”



- Assembly of charged particles (electrons & ions)
- Quasi-neutrality
- Debye shielding
- Interactions with electromagnetic field

# Rockets in plasma: What happens?

- The solid surface of a rocket introduces different boundary conditions for space plasma from its natural state  
⇒ Plasma density, flow velocity, etc. may change near the boundary.
- Plasma charge is deposited to its surface
  - The surface potential changes.  

  - The surface potential produces electric field.  

  - **Quasi-neutrality breaks down** near the surface.

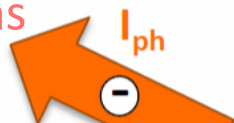


# Equilibrium potential

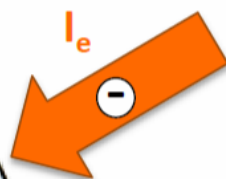
Charging is a result of the current balance:

$$\sum I = 0 \rightarrow I_e(V) - I_i(V) - I_{ph}(V) - I_{sec}(V) - I_{back}(V) = 0$$

Photo-  
electrons



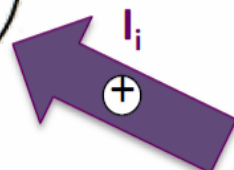
Plasma electron



$I_{sec}, I_{back}$

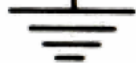


Secondary  
electrons



Plasma ions

$-V$



Space (reference) potential

$I_e$ : Net incoming electron current

$I_i$ : Net incoming ion current

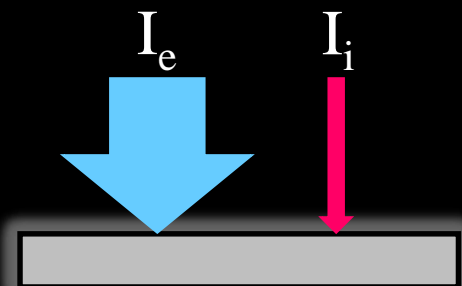
$I_{ph}$ : Net emitted photoelectron current

$I_{sec}$ : Net secondary electron current

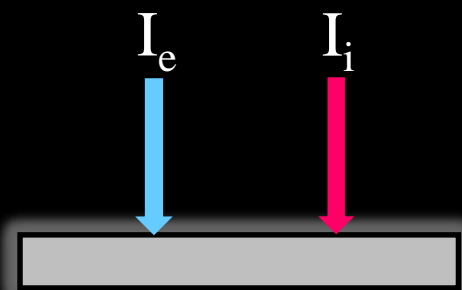
$I_{back}$ : Net backscattered electron current

# A few simple cases

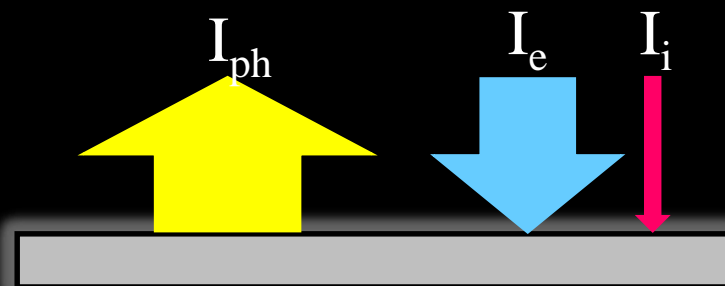
Case 1. Dense plasma consisting of only electrons & ions



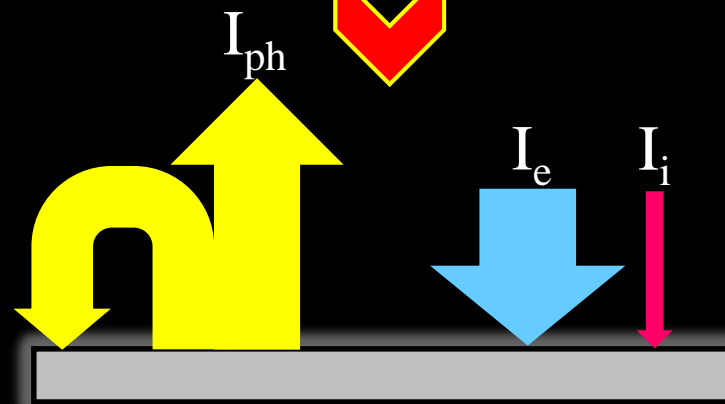
Potential goes DOWN



Case 2. Dilute plasma, in which photo-electron emission is NOT negligible



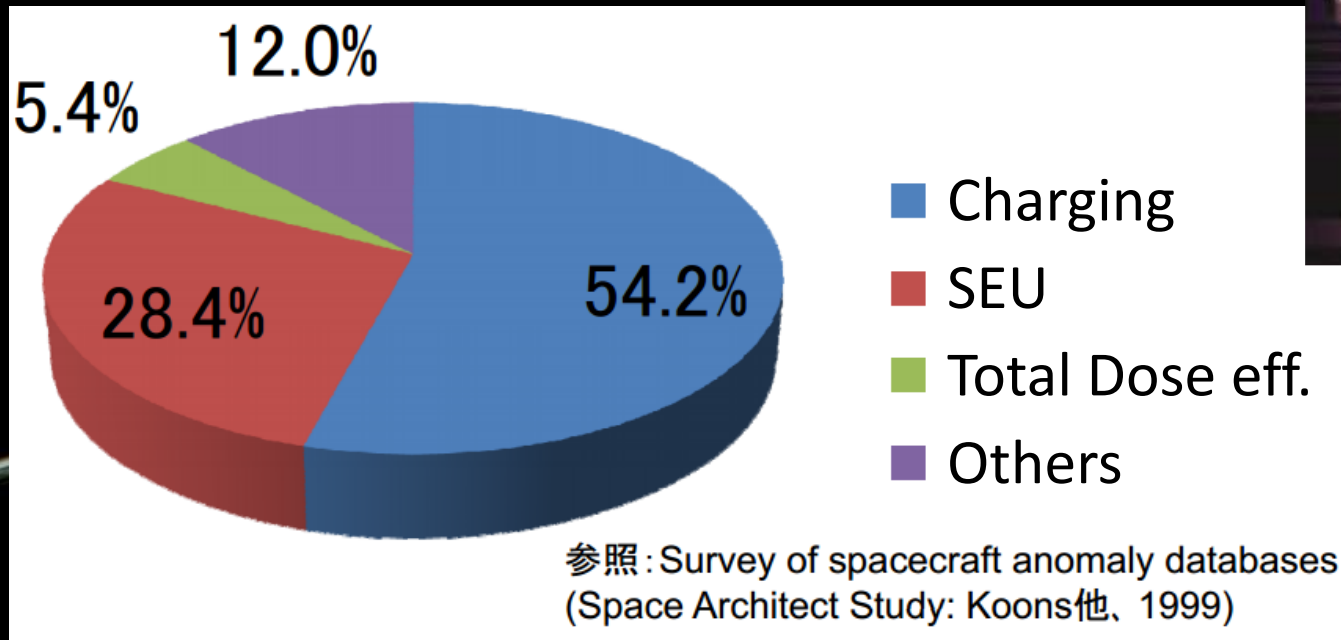
Potential goes UP



Current balance is established

# Why charging so important?

1. Charging/discharging may cause anomalies of in-space system.



2. Electromagnetic/plasma perturbation around rockets  
– Interference with scientific measurements (both fields and particles)

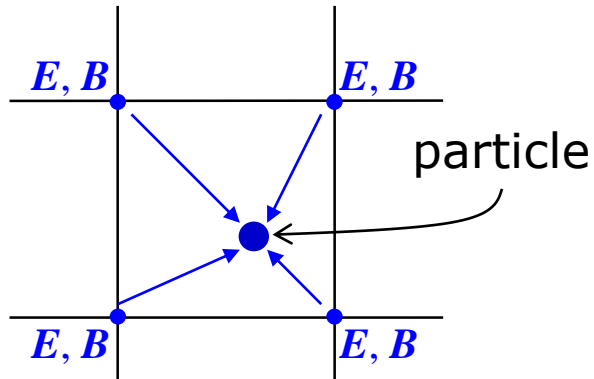
A space-themed background image showing Earth from space, a satellite in orbit, and a bright star in the upper left. The text is overlaid on this scene.

EMSES:

Electro-Magnetic Spacecraft Environment Simulator  
= *Full PIC code + Rocket body treatment*

EMSES is an electromagnetic particle-in-cell code designed for spacecraft-plasma interaction study. It can include spacecraft/rocket surfaces as an internal boundary, and has capability of simulating spacecraft charging, sheath/wake formation, etc., in a self-consistent manner.

# EMSES simulation: main loop



Field interpolation

$$\frac{d\mathbf{v}}{dt} = \frac{q_s}{m_s} (\mathbf{E} + \mathbf{v} \times \mathbf{B})$$
$$\frac{d\mathbf{r}}{dt} = \mathbf{v}$$

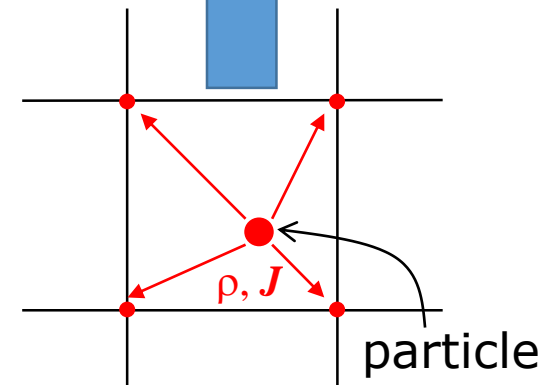
Particle dynamics

$$\nabla \times \mathbf{E} = -\frac{\partial \mathbf{B}}{\partial t}$$
$$\nabla \times \mathbf{B} = \frac{1}{c^2} \frac{\partial \mathbf{E}}{\partial t} + \mu_0 \mathbf{J}$$

or

$$\nabla \cdot (-\nabla \phi) = \frac{\rho}{\epsilon_0}$$

Field evolution

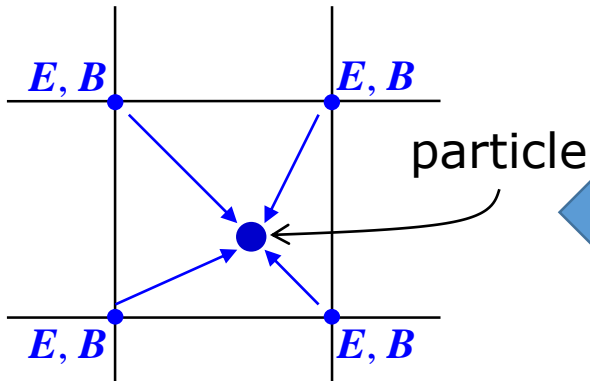


Current/charge deposition

$\Delta t$

# Mini-code: test-particle analysis

Using the field profile obtained from EMSES simulation



Field interpolation

$$\frac{d\mathbf{v}}{dt} = \frac{q_s}{m_s} (\mathbf{E} + \mathbf{v} \times \mathbf{B})$$
$$\frac{d\mathbf{r}}{dt} = \mathbf{v}$$

Particle dynamics

$\Delta t$

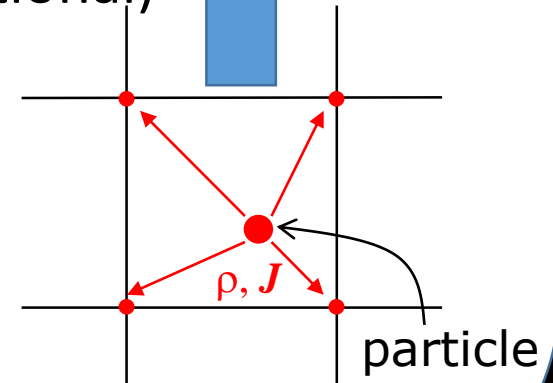
$$\nabla \times \mathbf{B} = \frac{1}{c^2} \frac{\partial \mathbf{E}}{\partial t} + \mu_0 \mathbf{J}$$

or

$$\nabla \cdot (-\nabla \phi) = \frac{\rho}{\epsilon_0}$$

Field evolution

(optional)



Current/charge deposition



# Test particle analysis

- Test particle analysis resolves the charged particle dynamics under the given electromagnetic field environment. (EM field is constant)
- Motion of each particle dynamics is solved individually.

## Test particle analysis

(Cons) ...does not include ANY plasma collective effects

(Pros) ...will be instructive to understand how particles are accelerated or decelerated in the given field

(Pros) ...will describe particle dynamics correctly under the assumption that the field environment reaches a steady state